Physics of the Electron Ion Collider: *Experimental* Introduction Part 2

Rosi Reed Lehigh University









Outline of Lectures

• Lecture 1: Particle Physics Detectors and Intro to ePIC

- Interaction of Particles with Matter
- Detector requirements at the EIC
- The design of the ePIC Detector
- Tracking Detectors

• Lecture 2: ePIC Detector and the EIC Science Case

- Calorimetry
- Particle Identification
- Far forward/Far backwards
- Physics observables
- Interaction Region Considerations

Particle Identification (PID) Detectors

- ePIC PID detectors are: dRICH, pfRICH, hpDIRC, AC-LGAD
- Other detector technologies also exist
- There are other ways in which the ID of a particle can be obtained using detectors that are not dedicated PID detectors
 - Electron ID (eID) can use the combination of tracking and calorimetry
 - dE/dx in detectors such as Time Projection Chambers (TPCs) can also give PID along with kinematic measurements



Time Projection Chamber (TPC)



TPCS use ionization drift time to determine the 3^{rd} coordinate of the "hit" (z = vt)

• Good for solenoidal B-fields where E||B Readout is at the and needs to be in a configuration that suppresses ion backflow

Charge buildup in the TPC will distort particle tracks!

6/9/2023



ALI-PERF-107348

Particle ID ePIC

PID needs

required detector length (cm) 00 00 00 00

10

- Electrons from photons $\rightarrow 4\pi$ coverage in tracking
- Electrons from charged hadrons → **mostly Calorimetry+tracking**
- Charged $\pi/K/p$ on track level \rightarrow **Cherenkov detectors**
 - Cherenkov detectors, complemented by ToF

Rapidity	π/K/p and πº/γ	e/h	Min p _T (E)	
-3.51.0	7 GeV/c	18 GeV/c	100 MeV/c	
-1.0 - 1.0	8-10 GeV/c	8 GeV/c	100 MeV/c	
1.0 - 3.5	50 GeV/c	20 GeV/c	100 MeV/c	



Need more than one technology to cover the entire momentum ranges at different rapidities

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Cerenkov Counters

 $\oint \cos \theta_C = \frac{1}{n\beta}$

- Cerenkov light emitted when v_{particle} > speed of light in a medium
 - Similar to a sonic boom (but optical)
 - Light emitted at a characteristic angle:
 - Light depends on particle velocity
 - Different momentum for different masses
 - Threshold Counter:
 - Just detect the light!
 - Ring Imaging Cerenkov (RICH):
 - Image the ring, measure the angle
 - Can be used to identify different particle types
 - Typically, only a few photons, measurements are delicate!



Cerenkov radiation from nuclear fission reactor core.

Cerenkov Counters

- Large, underground water Cerenkov counters were instrumental in demonstrating solar neutrino oscillations.
- Cerenkov light is used to detect cosmic rays in air showers









pfRICH – proximity focusing RICH

- Cherenkov radiator
 - 2.5 cm thick aerogel (n = 1.04-1.05)
 - 300 nm acrylic filter
 - $\langle N_{pe} \rangle \simeq 11-12$
 - Provide time with σ < 20 ps





IP

hpDIRC – High-Performance DIRC - ePIC

- DIRC (Detection of Internally Reflected Cherenkov light)
 - Uses total internal reflection of Cherenkov light within a solid radiator for precise particle identification
- ePIC hpDIRC components
 - 3-layer spherical lenses
 - Compact fused silica expansion volumes
 - Fast photodetection, small pixel MCP-PMT





dRICH – dual-radiator RICH - ePIC

- Radiators: aerogel (n \sim 1.02) and C2F6 (n \sim 1.0008)
- Mirrors: large outward-reflecting, 6 open sectors
- Sensors: 3x3 mm² pixel, 0.5 m² / sector
 - Single-photon detection inside high B field (> 1 T)



example event (accumulated hits)





Avalanche



 ΔV_{λ} = change in voltage due to avalanche

AC-LGAD Time of Flight (TOF) Detector

- Low-gain avalanche diodes (LGADs)
- AC-LGAD gain layer is not segmented, signal is read-out via AC coupling
 - Capacitive coupling → no direct electrical connection between the detector and the readout electronics, the signal is transferred through a thin insulating layer
- Charged particle generates electron-hole pairs
- High E-field in the LGAD layer amplifies signal (avalanche)
- Induces signal on the AC-coupled electrodes



AC-LGAD ePIC

- Two AC-LGAD layers
 - Barrel, |η| < 1.4
 - Forward, 1.5 < η < 3.5
- Barrel
 - 500 µm x 1 cm strips
 - 1% X₀
- Forward
 - 500 µm x 500 µm pixels
 - 8% X0
- Performance
 - Spatial resolution: 30 μm
 - Time resolution: $\sigma \sim 25 \text{ ps}$



AC-LGAD ePIC



Barrel Layer

- e/π up to 0.5 GeV
- π/K up to 1.9 GeV
- K/p up to 3.1 GeV

Forward Layer

- e/π up to 0.8 GeV
- π/K up to 2.7 GeV
- K/p up to 4.6 GeV



Particle ID ePIC

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TOF

1

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Scintillation Counters

- Basic idea is to convert the ionization to light, then collect the light
 - Has the advantage of a high efficiency even at low energies
- Organic Scintillators are typically a plastic (or oil) doped

with a fluorescing chemical.

- Passage of ionizing radiation excites the molecules and sets them into motion (vibration)
- Excited state spacings are of order ~eV, vibrational spacings around 1/10 eV
- Excited states decay to the vibrational ground state quickly (~ps), then later to the electronic ground state (~10ns)
- As a result, the material is essentially transparent to its own radiation!
 - Since most molecules are in the ground only one of many possible transition photons is likely to be absorbed.
 - The doping elements shift the UV light to visible



Figure 7.15 Electronic structure in an organic scintillator. The electronic states are represented as a potential minimum, resulting from the combined effects of the molecular attraction that keeps us from separating the atoms to greater distances and the repulsion that keeps us from forcing the atoms closer together (because the Pauli principle does not let the atomic wave functions overlap). Inside the electronic potential minimum is a sequence of levels that result from the atoms of the molecule vibrating against one another.



Reminder: Electromagnet Showers

Resolution of EM Calorimeters

$$\frac{\sigma_E}{E} \sim \frac{3\% - 10\%}{\sqrt{E/\text{GeV}}}.$$

- Number of EM particles essentially doubles every radiation length
- Light is collected to determine energy
- Shower develops until $\rm E_{ave}$ falls below the critical energy $\rm E_{c}$
 - Electrons and positrons then lose energy by ionization



Reminder: Hadronic Showers



Resolution of Hadronic Calorimeters

$$\frac{\sigma_E}{E} \gtrsim \frac{50\%}{\sqrt{E/\text{GeV}}}$$

material	X ₀ (g/cm ²)	λ_n (g/cm ²)	
H ₂	63	52.4	
AI	24	106	
Fe	13.8	132	
РЬ	6.3	193	

- Hadronic showers are messier and larger than EM showers
- Electromagnetic component due to $\pi^0 \rightarrow \gamma\gamma$
- Nuclear interaction interaction length λ_n → mean distance between hadronic interactions
 - Much larger than radiation length!

Energy Resolution in Calorimetry

$$\frac{\sigma}{E} = \frac{a}{\sqrt{E}} \bigoplus \frac{b}{E} \bigoplus C$$

Stochastic term (a): Fluctuations in the number of signal generating processes, for example the number of electrons that are detected as signals

Noise term (b): Noise in readout electronics, 'pile up' from other collision events close in time, etc.

Constant term (c): Imperfections in calorimeter construction, nonuniform response, calibration errors, energy lost in dead material (before or in detector)

A small constant term is crucial for good energy resolution at the highest particle energies.



$ePIC Detector \rightarrow EMCal$



EIC Detector Requirements EMCal

Electromagnetic calorimetry \rightarrow Measure photons (E, angle), identify electrons

- PbWO₄ Crystals (backward), W/ScFi (forward)
- Barrel Imaging Calorimeter (Si + Pb/ScFi)

Detailed EMCal Requirements

	Electron endcap (Backward)	Barrel	Hadron endcap (Forward)		
Energy Resolution	$\frac{(2-3)\%}{\sqrt{E}} \oplus 1\%$	$\frac{(10-12)\%}{\sqrt{E}}\oplus 2\%$	$\frac{(10-12)\%}{\sqrt{E}}\oplus 2\%$		
Shower Energy range	0.1–18 GeV	0.1–50 GeV	0.1–100 GeV		
$\pi \pm$ suppression (helped by other subsystems)	Up to 10 ⁴				
π^{0}/γ discrimination	Up to 18 GeV/c Up to 10 GeV/c		Up to 50 GeV/c		
Rad dose (includes background) at 10 ³⁴ sm ⁻² sec ⁻¹	<3 krad/year	<0.1 krad/year	<4 krad/year		
Max hit rate per tower (includes background)	50 kHz	5 kHz	20 kHz		
Neutron flux, at 10 ³⁴ sm ⁻² sec ⁻¹	10 ¹⁰ /cm ² /year	10 ¹⁰ /cm ² /year	10 ¹¹ /cm ² /year		
Limited space	Compact (small X ₀)				
Material on the way	Minimized				



Distinguishing between π^0 and γ is complicated for high momentum π^0 s due to cluster merging

ePIC Electromagnetic Calorimetry



EEEMCal

- Electron Endcap: PbWO₄
 - Key for finding DIS electron \rightarrow Defines DIS kinematics
 - High resolution, particularly at small angles
 - Strong hadron background suppression
 - Radiated γ measurements down to 100 MeV
 - DVCS photons (Deeply Virtual Compton Scattering)
 - Also decay photons (π^0/γ) and decay e











Hadron Endcap

W + SciFi SPACAL Design sPHENIX EMCal: 25k towers

- Compact: $X_0 = 0.7$ cm
- High granularity: Rm = 2cm
- Sampling fraction: ~2.3%
- Good resolution
- Jets up to energies ~100 GeV
- $\pi 0/\gamma$ discrimination to 50 GeV
- Decay electrons



Barrel: Pb/SciFi + Imaging

Hybrid Concept:

- 6 imaging Si layers (4 layers in baseline)
- Interleaved w/5 Pb/SciFi layers
- Followed by a thick Pb/SciFi layer (17X₀ total)
- Pb/SciFi → Scintillating fibers embedded in Pb (Similar to GlueX barrel EMCal)
- Imaging: Monolithic silicon sensor AstroPix (NASA's AMEGO-X mission)



Hadronic Calorimetry (sPHENIX→ ePIC)

Outer HCAL

Inner HCAL (Frame) MAGNET

EMCAL

(Frame $\approx 0.25\lambda_{I}$) (SPHENIX)

 $(EMCAL \approx 18X_0 \approx 0.7\lambda_1)$

Outer HCAL ≈3.5λ

Magnet ≈1.4X₀



- HCAL steel and scintillating tiles with wavelength shifting fiber
 - Outer HCal (outside the solenoid)
 - Δη x Δφ ≈ 0.1 x 0.1
 - 1,536 readout channels
- SiPM Readout

HCAL performance requirements driven by jet physics in HI collisions

- •Uniform fiducial acceptance -1< η <1 and 0< φ <2 π
 - Extended coverage -1.1< η <1.1 to account for jet cone
- (sPHENIX) Absorb >95% of energy from a 30 GeV jet
 - Requires ~4.9 nuclear interaction length depth
- (sPHENIX) Hadronic energy resolution of *combined* calorimetry:
 - UPP: $\frac{\sigma}{E} < \frac{150\%}{\sqrt{E}}$ (in central Au+Au collisions)
 - Gaussian response (limited tails)
- •HCAL created by instrumenting barrel magnetic flux return

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Outer HCAL Design



Far-Forward and Far-Backward Detectors



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A Few Other Experimental Concerns

Why a Crossing Angle?

- Brings focusing magnets close to IP
 → high luminosity
- Beam separation without separation dipoles
 reduced synchrotron radiation background
 But significant loss of luminosity!

Solution: Crab crossing

- Head-on collision geometry is restored by rotating the bunches before colliding ("crab crossing")
- Bunch rotation ("crabbing") is accomplished by transversely deflecting RF resonators ("crab cavities")
- Actual collision point moves laterally during
 bunch interaction
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Detector Services and Integration

Detectors require:

- Support frames
- Power
- Cooling
- Signals out
- ...

All of this represents dead, un-instrumented material in the path of the particles you are trying to detect!

Clever engineering and integration is required to minimize to optimize the capabilities of the combined detector.



Detector Backgrounds/Radiation

- Three types of background/radiation sources we need to worry about:
 - Primary collisions
 - Synchrotron radiation
 - Beam-gas induced
- All of these couple with the beam parameters
 - Divergence, energy spread, crossing angle, bunch length
- Why are these "bad"?
 - Detector occupancy
 - Additional hits in detectors (especially for the tracker vertexing layers)
 - Affects PID detectors due to scattering and secondary interactions
 - Detector lifetime (radiation damage)

Primary Collisions (photons and charged)

• Primary source of *ionizing radiation* and *low energy neutrons*



Radiation doses from gamma and charged particles fairly low across most of ePIC. Regions of highest dose near the beam line - can reach \sim 100 kRad.

Primary Collisions (neutrons)

• Primary source of *ionizing radiation* and *low energy neutrons*



1 MeV equivalent neutron fluence highest near the beamline -> implications for SiPMs.

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Synchrotron Radiation

107

106

105

10⁴ 10³

10²

Hits freq. [Hz]

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- Extensive simulations by of SR by accelerator background WG group
 - SynRad+ modeling software
 - Input:
 - 3D model of beampipe
 - Beam emittance, current Magnet locations and fields
 - Output:
 - Synchrotron Radiation Position, Flux, Energy, Direction
- Synchrotron Radiation Mitigation
 - Photon absorber configuration:
 - Wider beam pipe for
 - 13.5 σ clearance in x
 - 23 σ clearance in y
 - Beampipe material/structure
 - 2-5 μm Au coating
 - Sawtooth/ridge texture for photon absorption



Beam-Gas Consequences

rates in kHz	5x41 GeV	5x100 GeV	10x100 GeV	10x275 GeV	18x275 GeV
DIS ep	12.5 kHz	129 kHz	184 kHz	500 kHz	83 kHz
hadron beam gas	12.2kHz	22.0kHz	31.9kHz	32.6kHz	22.5kHz
electron beam gas	2181.97 kHz	2826.38 kHz	3177.25 kHz	3177.25 kHz	316.94 kHz

Electron Beam: $e + H^2 \rightarrow e' + \gamma + H^2$



Proton Beam: $p/A + H^2 \rightarrow stuff$

(also a significant contribution to low-energy neutrons in detector hall)



Final Thoughts

- In order to determine the answers to the fundamental structure of nuclear matter we need to understand the kinematics and flavor of all particles from our electron-ion collisions
- A large-scale optimization regarding cost, speed, resolution, robustness, technology interplay
- Pieces of ePIC already exist and are working well!
- Beam tests are ongoing \rightarrow Simulation is good but not sufficient
- Detector subsystems are essentially separate projects

Things I didn't cover

- Read-out electronics
- Data Acquisition (DAQ), Streaming Read-Out
- Calorimeter Clustering Routines, Tracking algorithms, Jet Finding
- Algorithms that combine information from a variety of detectors (Particle Flow is an example)
- Far forward/far backward regions
- Unfolding and other corrections for efficiency and resolution
- And probably a lot else... Many years of working on detectors...













SPHE

STAR